

## Letters to the Editors

### Radial Buckling Modes Near the Source of an Exponential Column\*

A set of data deleted from the published version (1) of the author's thesis (2) has proved quite valuable as an illustrative exercise for students of nuclear reactor engineering. The solution of the Helmholtz "wave" equation for a homogeneous cylindrical exponential column (3) with azimuthally independent source in the  $z = 0$  plane is

$$\phi(r, z) = \sum_{k=1}^{\infty} A_{0k} J_0(\alpha_{0k} r) \sinh \gamma_{0k}(h - z),$$

where the buckling eigenvalues are  $B^2_{0k} = \alpha^2_{0k} - \gamma^2_{0k}$  and the  $A_{0k}$  depend upon the source conditions.

TABLE I  
OBSERVED RELATIVE FLUX DISTRIBUTION

Radius (in.)	Observed total foil activation (counts/min)
0	$18.740 \times 10^4$
1	$17.785 \times 10^4$
2	$15.455 \times 10^4$
3	$13.161 \times 10^4$
4	$10.675 \times 10^4$
5	$8.297 \times 10^4$
6	$6.191 \times 10^4$
7	$4.575 \times 10^4$
8	$3.232 \times 10^4$
9	$2.188 \times 10^4$
10	$1.318 \times 10^4$

$$R(r) = 1.201^{\pm 0.025} \times 10^5 J_0 \left( \frac{2.405}{2.405} \cdot 0.2111^{\pm 0.0035} r \right) + 4.62^{\pm 0.25} \times 10^4 J_0 \left( \frac{5.520}{2.405} \cdot 0.2111^{\pm 0.0035} r \right) + 1.62^{\pm 0.16} \times 10^4 J_0 \left( \frac{8.654}{2.405} \cdot 0.2111^{\pm 0.0035} r \right)$$

Clearly, near a source of the column as in Figure 3, of ref. 1, the radial portion,  $R(r)$ , of the general solution should show the higher radial buckling modes,  $\alpha_{0k}$ ,  $k = 1, 2, 3, \dots$ . The boundary condition that  $R(\rho + d) = 0$  is satisfied by containing a series of  $J_0$  Bessel functions such that the  $k$ th zero of the  $k$ th term is at  $\rho + d$ , where  $d$  is the extrapolation distance to be determined.

The radial flux distribution in a natural uranium column

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with  $\rho = 10.5$  in., 4.1 in. from the source plane, was observed by foil activation techniques. The Los Alamos least-squares computer code (4) was used to separate the higher modes by fitting the function

$$R(r) = \sum_{k=1}^n b_k J_0(\alpha_{0k} r)$$

where the constraining boundary condition is met by

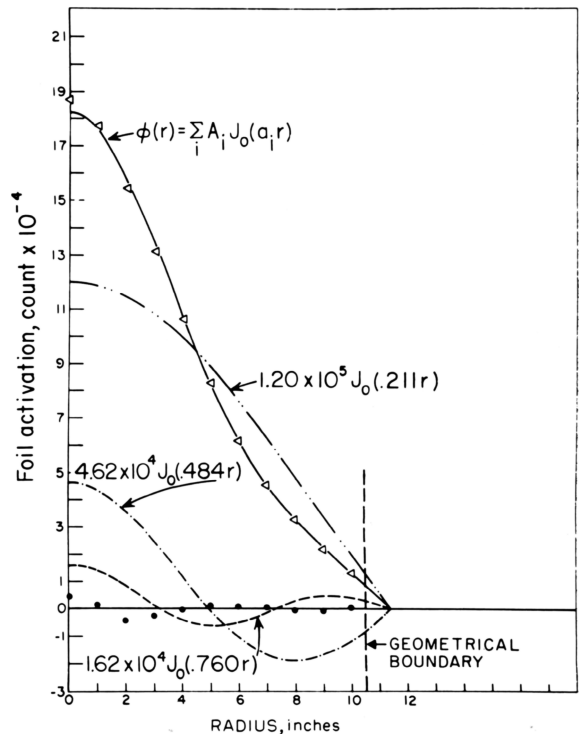


FIG. 1. Radial flux distribution, 4.1-in. level

setting

$$\alpha_{0k} = \frac{j_{0k}}{j_{01}} \alpha$$

leaving

$$\alpha = j_{01}/(\rho + d)$$

the only free parameter in the argument. The  $j_{0k}$  are the  $k$ th zeros of the  $J_0$  Bessel function.

Table I lists the typical corrected observations of  $\gamma$ -counts/min from the foils ( $U^{235}$ ,  $U^{238}$  or natural U) as well

as results of the least squares fit. Figure 1 graphically displays the three ( $n = 3$ ) modes determined by the computer. Note that in Fig. 1 the residuals are plotted as dots leading to the suspicion of the existence of another mode. The student should be able to justify, on a statistical basis, the inability of the computer program to resolve a fourth mode.

## REFERENCES

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CURTIS G. CHEZEM

*University of California  
Los Alamos Scientific Laboratory  
Los Alamos, New Mexico  
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